

27.2 Using Wien's displacement law,

$$(a) \quad \lambda_{\max} = \frac{0.2898 \times 10^{-2} \text{ m} \cdot \text{K}}{10^4 \text{ K}}$$

$$= 2.898 \times 10^{-7} \text{ m} \quad \boxed{\sim 100 \text{ nm}} \quad \boxed{\text{Ultraviolet}}$$

$$(b) \quad \lambda_{\max} = \frac{0.2898 \times 10^{-2} \text{ m} \cdot \text{K}}{10^7 \text{ K}} = 2.898 \times 10^{-10} \text{ m} \quad \boxed{\sim 10^{-1} \text{ nm}} \quad \boxed{\gamma\text{-rays}}$$

27.12 From Einstein's photoelectric effect equation,

$$e(\Delta V_s) = KE_{\max} = hf - \phi$$

$$\text{Thus, } \phi = hf - e(\Delta V_s) = (6.63 \times 10^{-34} \text{ J}\cdot\text{s})(3.0 \times 10^{15} \text{ Hz}) \left(\frac{1 \text{ eV}}{1.6 \times 10^{-19} \text{ eV}} \right) - e(7.0 \text{ V})$$

$$\text{or } \phi = 12.4 \text{ eV} - 7.0 \text{ eV} = \boxed{5.4 \text{ eV}}$$

27.13 From Einstein's photoelectric effect equation,

$$KE_{\max} = hf - \phi \quad \text{or} \quad KE_{\max} = \frac{hc}{\lambda} - \phi$$

$$\text{Thus, } \lambda = \frac{hc}{KE_{\max} + \phi} = \frac{hc}{\frac{1}{2}m_e v_{\max}^2 + \phi}$$

$$\text{or } \lambda = \frac{(6.63 \times 10^{-34} \text{ J}\cdot\text{s})(3.00 \times 10^8 \text{ m/s})}{\frac{1}{2}(9.11 \times 10^{-31} \text{ kg})(1.00 \times 10^6 \text{ m/s})^2 + (2.46 \text{ eV})(1.60 \times 10^{-19} \text{ J/eV})}$$

$$\lambda = 2.34 \times 10^{-7} \text{ m} = \boxed{234 \text{ nm}}$$

27.39 For relativistic particles, $p = \frac{\sqrt{E^2 - E_R^2}}{c}$ and $\lambda = \frac{h}{p} = \frac{hc}{\sqrt{E^2 - E_R^2}}$

For 3.00 MeV electrons, $E = KE + E_R = 3.00 \text{ MeV} + 0.511 \text{ MeV} = 3.51 \text{ MeV}$

so
$$\lambda = \frac{(6.63 \times 10^{-34} \text{ J} \cdot \text{s})(3.00 \times 10^8 \text{ m/s})}{\sqrt{(3.51 \text{ MeV})^2 - (0.511 \text{ MeV})^2}} \left(\frac{1 \text{ MeV}}{1.60 \times 10^{-13} \text{ J}} \right) = \boxed{3.58 \times 10^{-13} \text{ m}}$$